# Understanding Exoplanet Habitability: An AI framework for Predicting Atmospheric Absorption Spectra

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## Abstract

Evolution of space technology, fueled by advancements in Artificial Intelligence and Machine Learning, has transformed our capacity to explore space. The James Webb Space Telescope (JWST) has made this information more accessible, resulting in valuable data. We are working to create an atmospheric absorption spectrum prediction model for exoplanets based on spectral data generated by General Circulation Models developed by NASA's Goddard Institute for Space Studies. Describing an atmospheric absorption spectrum as a set of bins, a model of the spectral bin heights can be generated as a function of the estimated 30 planetary parameters. Spline curves describe the bin heights as a function of the values of the planetary parameters. The main difficulty is the dimensionality of the parameter space, which will include up to 30 planetary parameters. Once the spectrum prediction model is created, Bayesian Adaptive Exploration will identify areas of the planetary parameter space for which more data are needed to constrain the model better. Specifically, parameter values for which the uncertainties in the predicted spectra are largest (highest entropy) are candidates for generating synthetic spectra to improve the system's precision. As more synthetic spectra are generated to train the system, the model's predictive capabilities will increase. This system will be used as a forward model to assist in forming likelihood functions so that the planetary parameters can be inferred given a planet's atmospheric absorption spectrum. This endeavor is expected to contribute to a better understanding of exoplanetary properties and general habitability.

### Basic Idea

Basic idea is to utilize Measured Synthetic Atmospheric and Spectra sourced from NASA's Climate Models (GCMs) as training data for an AI system. This AI system will the work as a forward model to predict spectra based on a set of approximately 30 planetary parameters.

Further helping us understand the climatic conditions of any planet



- Develop an
- from spectra parameters.
- learning system to planetary parameters spectral data.



The spectra in image is a sample of collected data. The spectrum on the left comes from the ROCKE3D simulation. Leftmost being the visible range (VIS) followed by Near InfraRed bands.





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Methodology

# Logarithmic interpolation

On previous version,

- Implemented logarith interpolation formula
- Calculated logarithmic values splined value
- Visualize the curves

Mathematical Formula :

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**Optimization Goals:** 







References Way, M., Ackerman, A., Aleinov, I., Barnes, R., Chandler, M., Colose, C., Fauchez, T., Guzewich, S., Harman, "S".C., Kiang, N., Knuth, K.H., Leboissetier, A. Scharf, C., Sohl, L., Tsigaridis, K., Villanueva, G., Wolf, E. (2023). Habitability Space: Exploring a New Frontier via Climate Models and Planetary Statistics, [ A grant proposal to the NASA Interdisciplinary Consortia for Astrobiology Research].

	Nested Sampling
	On addition to previous version,
hmic	• Nested sampling is used as a Bayesian inference technique.
for	<ul> <li>Iteratively evaluate</li> </ul>
	 <ul> <li>Implement likelihood function</li> </ul>
	<ul> <li>Narrowed down parameter space</li> </ul>
	• Used Bayesian approach to refine
	parameter
	Mathematical Formula :
	$L( heta \mid D,M) = P(D \mid  heta,M)$

# Discoussion

# **Current Prototype Limitations**

• Limited to one/two-dimensional analysis with five spline knots per dimension (25 splines in two dimensions).

# High Dimensionality Challenge:

• Scaling to 30 dimensions would require impractically large numbers of splines (10<sup>2</sup>1), due to unaccounted correlations between spline knots.

• Reduce model parameters from billions to thousands by factorization.